

Stable optical spring in aLIGO detector and in dual-recycling Michelson-Sagnac interferometer

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INVESTMENTS IN EDUCATION DEVELOPMENT

Olomouc, 2014



Gravitational waves

Detectors

Standard Quantum Limit

Optical spring effect

Stable optical spring



Gravitational waves

Detectors

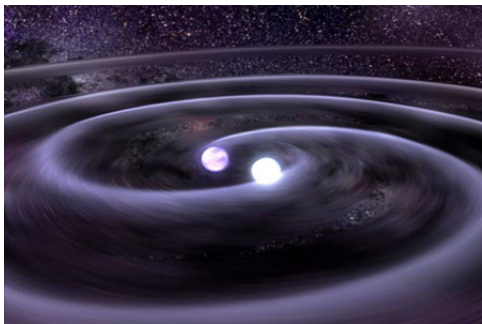
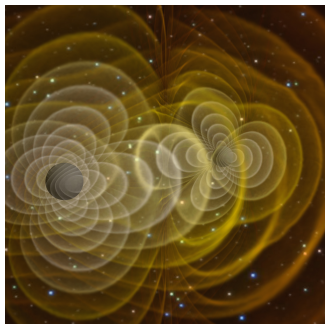
Standard Quantum Limit

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Gravitational waves

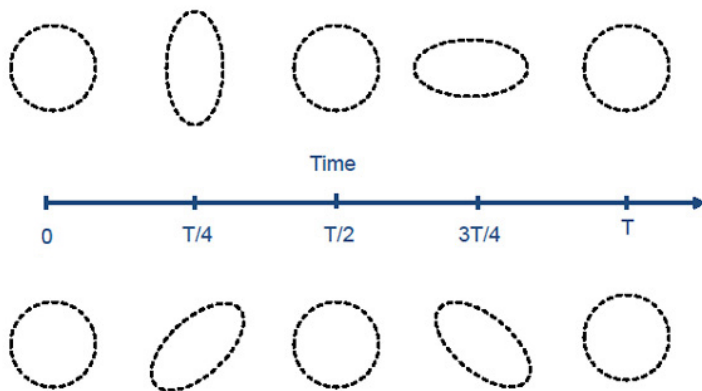


- 1) The new window to explore the Universe
- 2) The confirmation of the General Theory of Relativity

1



How do they interact with matter?



2



Gravitational waves

Detectors

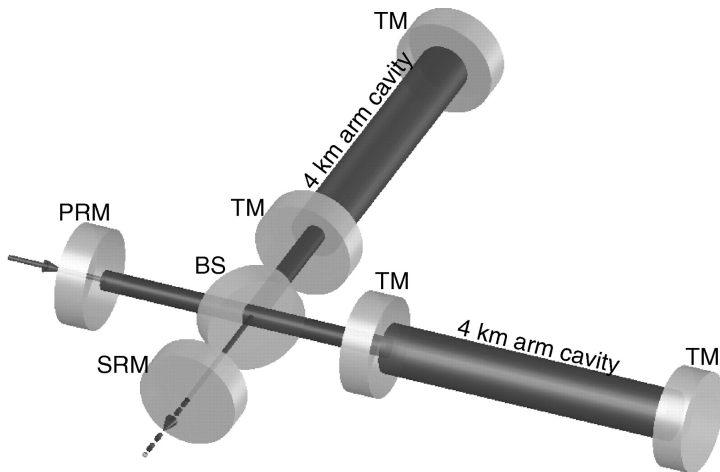
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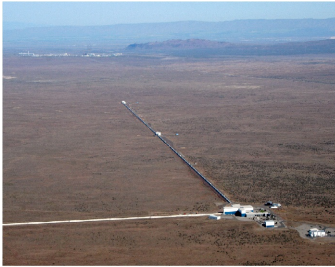
How can we catch them? Advanced LIGO



3



How can we catch them?



LIGO interferometers



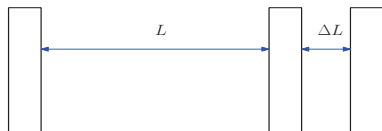
GEO 600



VIRGO



Order of effect



$$\Delta L = \frac{1}{2}hL, \quad L = 4km$$

► Initial LIGO:

$$h \sim 10^{-21} \Rightarrow \Delta L \sim 10^{-18}cm \quad \sim 0.5events/year$$

► Advanced LIGO:

$$h \sim 10^{-22} \Rightarrow \Delta L \sim 10^{-19}cm \quad \sim 10^3events/year$$

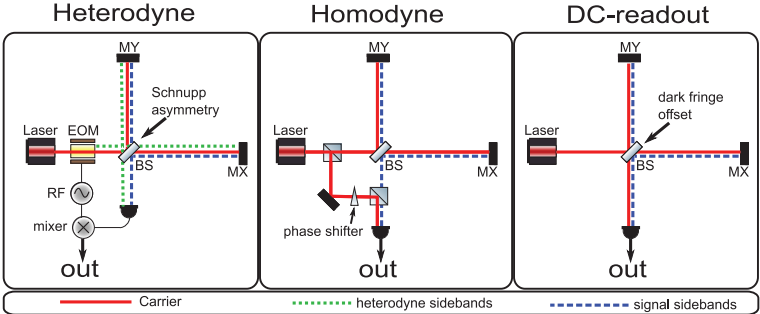


Possible detection regimes:

- ▶ Dark port regime
Arms of interferometer are balanced \rightarrow no mean power in the signal port
- ▶ Operating off-dark port
Arms of interferometer are unbalanced \rightarrow mean power in the signal port appears



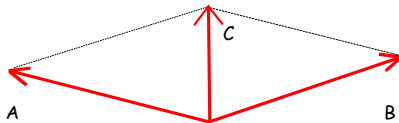
DC readout



Operating off-dark port



a) Balanced scheme



b) Unbalanced scheme



Gravitational waves

Detectors

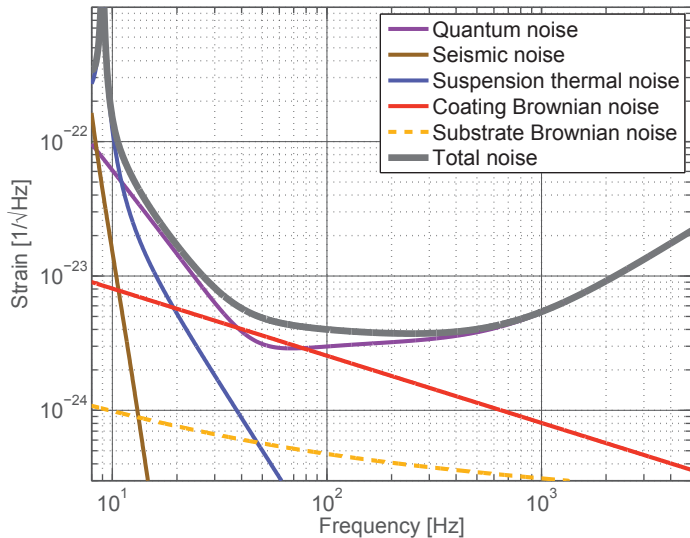
Standard Quantum Limit

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Noise budget of Advanced LIGO detector



Standard quantum limit

Quantum noise

The light produces two types of noise:

- ▶ Photon shot-noise
- ▶ Radiation pressure noise

$$S_h^{shot} \sim \frac{1}{\mathcal{I}_0}, \quad S_h^{rad} \sim \mathcal{I}_0$$

The optimal case:

$$S_h^{shot} = S_h^{rad}$$

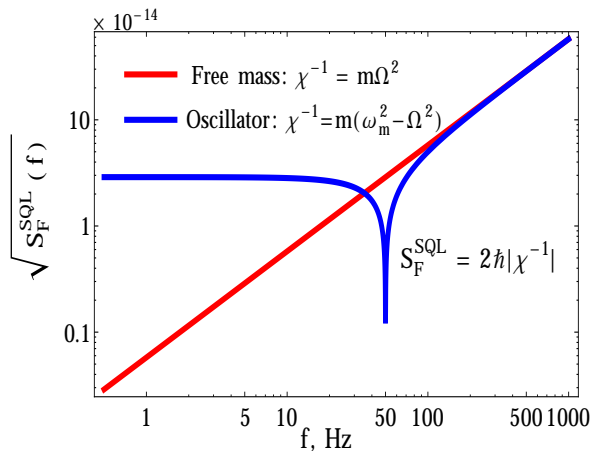
The resulting net noise:

$$S_h = S_h^{shot} + S_h^{rad} = 2S_h^{shot}$$



Standard quantum limit

This minimum achievable noise is called the “standard quantum limit” (SQL) and is denoted S_F^{SQL}



Gravitational waves

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Optical spring effect

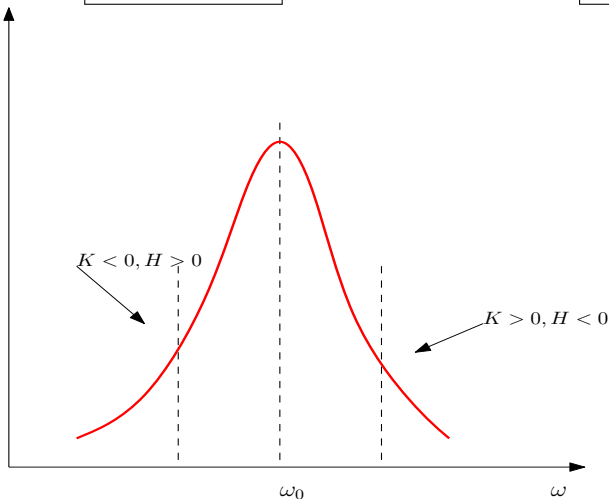
Stable optical spring



Optical spring effect

$$\omega_p - \omega_0 = \delta < 0 \Rightarrow K = -\frac{dF}{dx} < 0,$$

$$\omega_p - \omega_0 = \delta > 0 \Rightarrow K = -\frac{dF}{dx} > 0$$



Optical spring effect

- Circumvent of the SQL: The optical spring turns free mass into the oscillator and that allows to beat the SQL in some frequency band ⁶.
- Instability:

$$m_{eq}\ddot{y} + h\dot{y} + (K + \Delta K)y = 0,$$

$$y(t - \tau^*) \simeq y(t) - \tau^* \dot{y}(t)$$

$$m_{eq}\ddot{y}(t) + (h \mp \tau^* \Delta K) \dot{y}(t) + (K \pm \Delta K)y(t) = 0.$$

⁶V.B. Braginsky, F.Ya. Khalili, Physics Letters A, **257**, 241 (1999).
F.Ya. Khalili, PLA, **288**, 251-256 (2001);
A. Buonanno, Y. Chen, PRD, **64**, 042006 (2001);
V. I. Lazebny, S. P. Vyatchanin, PLA, **344**, 7 (2005);
F. Ya. Khalili, V. I. Lazebny, S. P. Vyatchanin, PRD **73**, 062002 (2006)



Ways to avoid instability

- 1) Feedback ⁷
- 2) Double optical spring ⁸

⁷A. Buonanno, Y. Chen, PRD, **65**, 042001 (2002)

⁸H. Rehbein, H. Müller-Ebhardt, S.L. Danilishin, C. Li, R. Schnabel, K. Danzmann, and Y. Chen, Phys. Rev. **D78**,062003 (2008),
A.A. Rakhubovsky, S. Hild, S.P. Vyatchanin, Phys. Rev. **D84**,062002 (2011)



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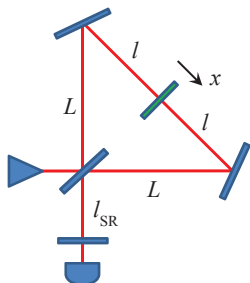
Stable optical spring



The Idea

It was shown that it is possible to achieve stable optical rigidity in interferometer operated off-dark port ⁹

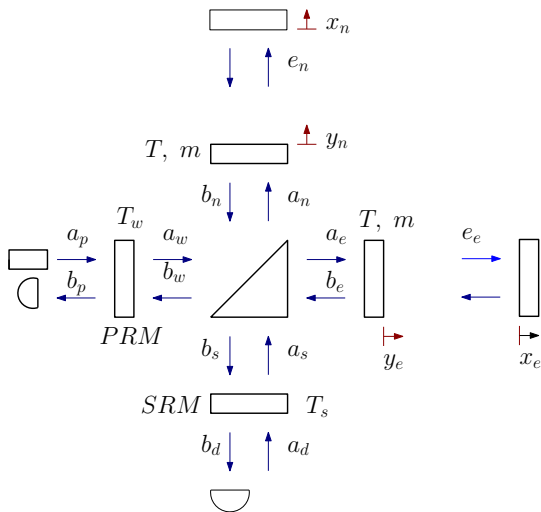
- ▶ It was shown on the example of the Michelson-Sagnac interferometer without power-recycling mirror.
- ▶ Relatively big arms detuning (operating far from dark port).



⁹S. P. Tarabrin, H. Kaufer, F. Ya. Khalili, R. Schnabel, and K. Hammerer, Physical Review A 88, 023809 (2013)



Advanced LIGO scheme



Some basic points

Dynamics of interferometer is described by two mechanical and two optical modes.

$$z_{e,n} = x_{e,n} - y_{e,n}$$

► Mechanical modes:

- 1) Differential mode (antisymmetric): $z_- = \frac{z_e - z_n}{2}$
- 2) Common mode (symmetric): $z_+ = \frac{z_e + z_n}{2}$

► Optical modes:

- 1) Differential mode

This mode is characterized by optical detuning δ_s and decay rate γ_s

- 2) Common mode

This mode is characterized by optical detuning δ_w and decay rate γ_w



Our aim

- ▶ To get the stability with single pump
- ▶ To find a minimal value of arms detuning necessary to obtain stable optical spring

Consideration and assumptions:

- ▶ Common and differential modes became weakly coupled
- ▶ Small decay rates and detuning in arms.

$$\gamma_w \ll \delta_w, \quad \gamma_s \ll \delta_s, \quad \delta \ll \delta_s, \quad \delta_w$$

- ▶ We exclude symmetric mechanical mode

$$z_+ \equiv \frac{(x_e - y_e) + (x_n - y_n)}{2} = 0$$



Initial equations

$$\left\{ \begin{array}{l} (\gamma_w - i\delta_w - i\Omega) e_+(\Omega) - i\delta e_-(\Omega) - \frac{ik}{\tau} E_{-z_-}(\Omega) = \frac{\sqrt{\gamma_w} g_p(\Omega)}{\sqrt{\tau}}, \\ -i\delta e_+(\Omega) + (\gamma_s - i\delta_s - i\Omega) e_-(\Omega) - \frac{ik}{\tau} E_{+z_-}(\Omega) = \frac{\sqrt{\gamma_s} g_d(\Omega)}{\sqrt{\tau}}, \\ (\gamma_w + i\delta_w - i\Omega) e_+^\dagger(-\Omega) + i\delta e_-^\dagger(-\Omega) + \frac{ik}{\tau} E_{-z_-}^*(\Omega) = \frac{\sqrt{\gamma_w} g_p^\dagger(-\Omega)}{\sqrt{\tau}}, \\ i\delta e_+^\dagger(-\Omega) + (\gamma_s + i\delta_s - i\Omega) e_-^\dagger(-\Omega) + \frac{ik}{\tau} E_{+z_-}^*(\Omega) = \frac{\sqrt{\gamma_s} g_d^\dagger(-\Omega)}{\sqrt{\tau}}, \\ \hbar k \{ E_+^* e_-(\Omega) + E_-^* e_+(\Omega) + E_+ e_-^\dagger(-\Omega) + E_- e_+^\dagger(-\Omega) \} + \mu \Omega^2 z_-(\Omega) = 0. \end{array} \right.$$

Notations



Normal coordinates

- Balanced aLIGO interferometer:

Common (e_+) and differential (e_-) modes:

$$e_+ = \frac{e_e + e_n}{\sqrt{2}}, \quad e_- = \frac{e_e - e_n}{\sqrt{2}}$$

- Unbalanced aLIGO interferometer:

Normal modes (b_{\pm}):

$$b_+ = e_+ - \kappa e_-, \quad b_- = \kappa e_+ + e_-, \quad \kappa \sim \delta$$

Characteristic equation:

$$\lambda^2 + \frac{\mathcal{I}_1 [1 + \alpha_1 (\lambda + \tilde{\gamma}_s)]}{(\lambda + \tilde{\gamma}_s)^2 + \tilde{\delta}_s^2} + \frac{\mathcal{I}_2 [1 + \alpha_2 (\lambda + \tilde{\gamma}_w)]}{(\lambda + \tilde{\gamma}_w)^2 + \tilde{\delta}_w^2} = 0$$



We can rewrite characteristic equation in the next form:

$$\underbrace{D_1^{(0)} D_2^{(0)}}_{\text{zero order}} + \underbrace{D^{(1)}}_{\text{first order} \sim \delta^2} = 0,$$

$$D_1^{(0)} = \lambda^2 [(\lambda + \tilde{\gamma}_s)^2 + \tilde{\delta}_s^2] + \mathcal{I}_1 [1 + \alpha_1(\lambda + \tilde{\gamma}_s)],$$

$$D_2^{(0)} = (\lambda + \tilde{\gamma}_w)^2 + \tilde{\delta}_w^2,$$

$$D^{(1)} = [(\lambda + \tilde{\gamma}_s)^2 + \tilde{\delta}_s^2] \mathcal{I}_2 (1 + \alpha_2(\lambda + \tilde{\gamma}_w)).$$

Zero order:

$$\underbrace{D_1^{(0)}}_{\lambda_{1,2}^{(0)}, \lambda_{3,4}^{(0)}} \times \underbrace{D_2^{(0)}}_{\lambda_{5,6}^{(0)}} = 0, \quad \lambda_{1,2}^{(0)} = \gamma_1 \pm i\delta_1, \quad \gamma_1 > 0 \text{ — instability,}$$

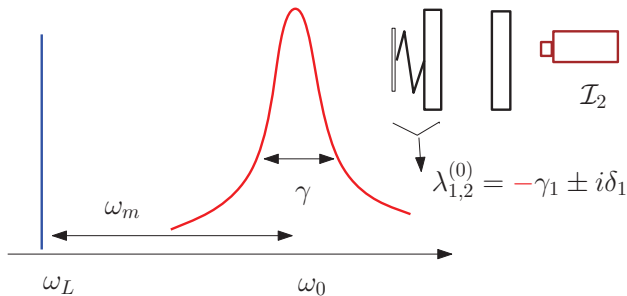
$$\lambda_{3,4}^{(0)} = -\gamma_3 \pm i\delta_3, \quad \lambda_{5,6}^{(0)} = -\tilde{\gamma}_w \pm i\tilde{\delta}_w, \quad \gamma_3, \tilde{\gamma}_w > 0$$



First order iteration

Problem: To introduce **maximal** relaxation into unstable mode by using “additional” pump \mathcal{I}_2 as small as possible.

Solution: We use an analogy of laser-cooling : $\tilde{\delta}_w = -\delta_1$



First order iteration

We choose $\tilde{\delta}_w = -\delta_1$

$$((\lambda - \gamma_1)^2 + \delta_1^2) ((\gamma + \tilde{\gamma}_w)^2 + \delta_w^2) - b = 0,$$

where

$$b \equiv - \left. \frac{D^{(1)}}{(\lambda + \gamma_3)^2 + \delta_3^2} \right|_{\lambda = \lambda_{1,2}^{(0)}}$$

We consider b as a constant of first order of smallness

$$\lambda = \frac{\gamma_1 - \tilde{\gamma}_w}{2} \pm i \sqrt{\delta_1^2 - \left[\frac{\gamma_1 + \tilde{\gamma}_w}{2} \right]^2} \pm \sqrt{b - \delta_1^2 [\gamma_1 + \tilde{\gamma}_w]^2}$$



Stability conditions

$$\lambda = \frac{\gamma_1 - \tilde{\gamma}_w}{2} \pm i \sqrt{\delta_1^2 - \left[\frac{\gamma_1 + \tilde{\gamma}_w}{2} \right]^2} \pm \sqrt{b - \delta_1^2 [\gamma_1 + \tilde{\gamma}_w]^2}$$

Stability conditions:

$$\Re b = \delta_1^2 [\gamma_1 + \tilde{\gamma}_w]^2, \quad \gamma_1 < \tilde{\gamma}_w$$

Estimation of minimal value of δ :

$$\delta \sim \tilde{\gamma}_w$$

$\tilde{\gamma}_w$ - decay rate of common optical mode.

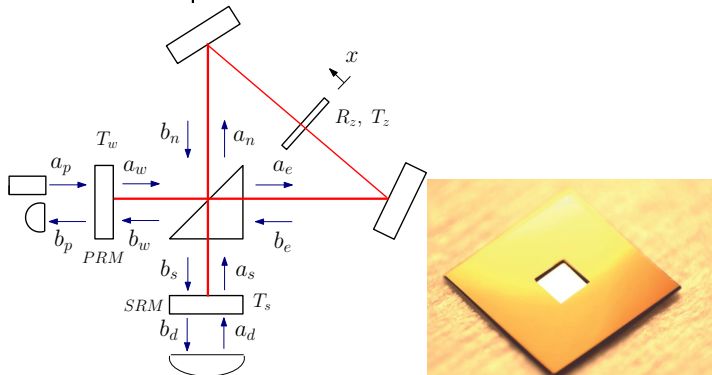


Michelson-Sagnac interferometer

We show that after substitution

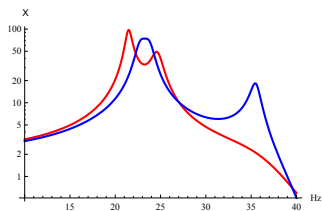
$$\delta^2 \rightarrow R_z^2 \delta^2, \quad J_+ \rightarrow R_z^2 J_+, \quad \mu \rightarrow m,$$

the characteristic equation is the same that for MSI.

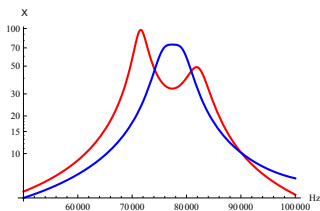


Numerical estimations

aLIGO



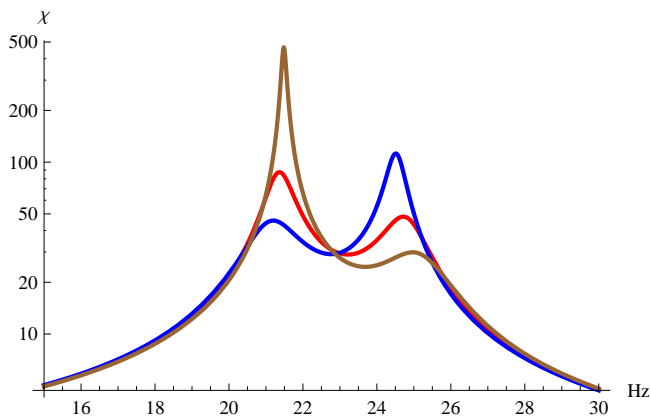
MSI



Common mode detuning (δ_w)	-23.0 Hz
Differential mode detuning (δ_s)	42.4 Hz
Decay rate of common mode (γ_w)	3.0 Hz (1.5)
Decay rate of differential mode (γ_s)	3.0 Hz (0.3)
Test mass (m)	40 kg
Arm length (L)	4 km
Circulating power (I_{circ})	24 kW
Arms detuning (δ)	4.6 Hz (1.5)
Output power (I_{out})	0.03 W

Common mode detuning (δ_w)	-77.2 kHz
Differential mode detuning (δ_s)	141.0 kHz
Decay rate of common mode (γ_w)	10 kHz (5)
Decay rate of differential mode (γ_s)	10 kHz (1)
Mass of membrane (m)	10^{-10} kg
Arm length (L)	8.7 cm
Circulating power (I_{circ})	318 mW
Arms detuning (δ)	15 kHz (5)
Membrane's reflectivity (R_z^2)	0.17

Shape of susceptibility curve



Susceptibility χ of a LIGO with $\tilde{\delta}_w = -\delta_1 + \Delta$. Red curve - $\Delta = 0$.
Blue curve - $\Delta = 0.5$ Hz. Brown curve - $\Delta = -0.5$ Hz



Summary

- ▶ Very high-sensitive detectors are required to detect GW.
- ▶ Quantum effects come into play. The SQL.
- ▶ Optical spring is one of the possible ways to beat SQL. Optical spring is unstable.
- ▶ It was shown that it is possible to achieve stable optical spring in the UNBALANCED detector with SINGLE PUMP.
- ▶ The stable optical spring may be created with small arm detuning comparable with optical bandwidths:

$$\delta \sim \tilde{\gamma}_{w,s}$$

- ▶ These results can be applied to the table top Michelson-Sagnac interferometer with membrane inside to create stable optical spring (and to another optomechanical systems)



Thank you!



Routh-Hurwitz stability criterion

$$D = \sum_{n=0}^6 a_n \lambda^n$$

Hurwitz matrices $H_1, H_2, H_3, H_4, H_5, H_6$.

Determinants of all matrices must be positive: $\text{Det}(H_i) > 0$.

Our analysis shows:

$$\begin{aligned} \text{Det}(H_1) &= 3.6, & \text{Det}(H_2) &= 2721.22, & \text{Det}(H_3) &= 1.4 * 10^6, \\ \text{Det}(H_4) &= 9.7 * 10^{11}, & \text{Det}(H_5) &= 2.9 * 10^{15}, & \text{Det}(H_6) &= 1.0 * 10^{24}. \end{aligned}$$

Hurwitz matrices



Differential and common optical modes:

$$e_{\pm} = \frac{e_e \pm e_n}{\sqrt{2}}, \quad E_{\pm} = \frac{E_e \pm E_n}{\sqrt{2}}.$$

We introduce a new basis:

$$g_p = e^{i\alpha_w} a_p, \quad g_d = e^{i\alpha_s} a_d.$$

so that

$$\begin{aligned} [g_d(\Omega), g_d^{\dagger}(\Omega')] &= 2\pi \delta(\Omega - \Omega'), \\ [g_p(\Omega), g_p^{\dagger}(\Omega')] &= 2\pi \delta(\Omega - \Omega'). \end{aligned}$$



Notations

Initial equations in normal coordinates:

$$\left\{ \begin{array}{l} (-i\Omega - \lambda_+) b_+(\Omega) - iz_- [\xi - \varkappa] = 0, \\ (-i\Omega - \lambda_-) b_-(\Omega) - iz_- [1 + \varkappa\xi] = 0, \\ (-i\Omega - \lambda_+^*) b_+^\dagger(-\Omega) + iz_- [\xi^* - \varkappa^*] = 0, \\ (-i\Omega - \lambda_-^*) b_-^\dagger(-\Omega) - iz_- [1 + \varkappa^* \xi^*] = 0, \\ \frac{b_+(\Omega) [\xi^* - \varkappa] + b_-(\Omega) [1 + \xi^* \varkappa]}{d} + \\ + \frac{b_+^\dagger(-\Omega) [\xi + \varkappa^*] + b_-^\dagger(-\Omega) [1 + \xi \varkappa^*]}{d^*} + \frac{\Omega^2}{J_+} z_- = 0, \end{array} \right.$$



Notations

$$\lambda_{\pm} = -\left(\Gamma_{+} \pm \Gamma_{-} \sqrt{1 + \Delta^2}\right), \quad J_{+} \equiv \frac{kl_{+}}{L\mu},$$

$$\Gamma_{\pm} \equiv \frac{\gamma_w - i\delta_w \pm (\gamma_s - i\delta_s)}{2}, \quad d \equiv 1 + \varkappa^2,$$

$$\varkappa \equiv \frac{i\delta}{\Gamma_w + \lambda_{-}} = \frac{\Delta}{1 + \sqrt{1 + \Delta^2}}, \quad \xi \equiv \frac{i\delta}{\gamma_s - i\delta_s}, \quad \Delta \equiv \frac{i\delta}{\Gamma_{-}}.$$

We introduce

$$\lambda_{+} \equiv -\tilde{\gamma}_w + \tilde{\delta}_w, \quad \lambda_{-} \equiv -\tilde{\gamma}_s + \tilde{\delta}_s,$$

$$\mathcal{I}_1 \equiv \frac{2J_{+}\tilde{\delta}_s \Re\phi}{|d|^2}, \quad \alpha_1 \equiv \frac{\Im\phi}{\tilde{\delta}_s \Re\phi}, \quad \mathcal{I}_2 \equiv \frac{2J_{+}\tilde{\delta}_w \Re\psi}{|d|^2}, \quad \alpha_2 \equiv \frac{\Im\psi}{\tilde{\delta}_w \Re\psi},$$

$$\phi \equiv (1 + \xi^* \varkappa)(1 + \varkappa \xi)d^*, \quad \psi \equiv (\xi^* - \varkappa)(\xi - \varkappa)d^*.$$



Hurwitz matrices

$$H_1 = a_1,$$

$$H_2 = \begin{pmatrix} a_1 & 1 \\ a_3 & a_2 \end{pmatrix}, \quad H_3 = \begin{pmatrix} a_1 & 1 & 0 \\ a_3 & a_2 & a_1 \\ a_5 & a_4 & a_3 \end{pmatrix}, \quad H_4 = \begin{pmatrix} a_1 & 1 & 0 & 0 \\ a_3 & a_2 & a_1 & 1 \\ a_5 & a_4 & a_3 & a_2 \\ a_7 & a_6 & a_5 & a_4 \end{pmatrix},$$

$$H_5 = \begin{pmatrix} a_1 & 1 & 0 & 0 & 0 \\ a_3 & a_2 & a_1 & 1 & 0 \\ a_5 & a_4 & a_3 & a_2 & a_1 \\ a_7 & a_6 & a_5 & a_4 & a_3 \\ 0 & a_8 & a_7 & a_6 & a_5 \end{pmatrix}, \quad H_6 = \begin{pmatrix} a_1 & 1 & 0 & 0 & 0 & 0 \\ a_3 & a_2 & a_1 & 1 & 0 & 0 \\ a_5 & a_4 & a_3 & a_2 & a_1 & 1 \\ a_7 & a_6 & a_5 & a_4 & a_3 & a_2 \\ 0 & a_8 & a_7 & a_6 & a_5 & a_4 \\ 0 & 0 & 0 & a_8 & a_7 & a_6 \end{pmatrix}$$

Back

