

Theoretical Quantum Optics

Gaussian entanglement and beyond in atmospheric channels

Martin Bohmann

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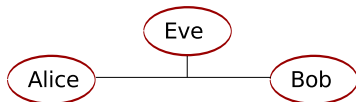
J. Sperling
University of Oxford

Why quantum optics in the atmosphere?

The question can be answered in two parts:

1. **Quantum part:**

- Secure data communication with the aid of quantum physics
- Quantum key distribution

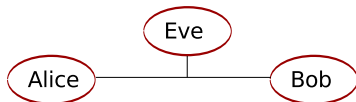


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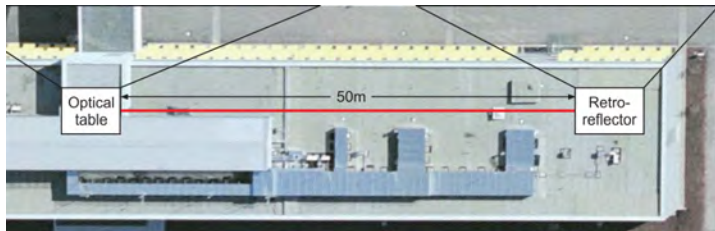
- Secure data communication with the aid of quantum physics
- Quantum key distribution



2. Atmosphere part:

- Optical fibers are limited due to losses ($\sim 100\text{km}$)
- Flexible global communication possible

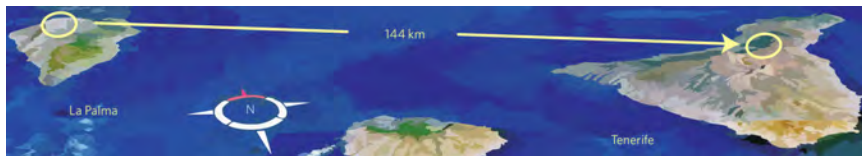
From rooftops to satellites



| [references](#)

D. Elser *et al.*, *New J. Phys.* **11**, 045014 (2009) [1]

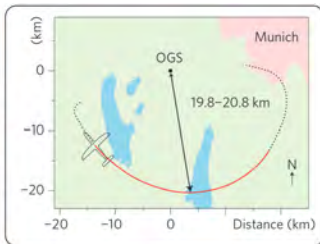
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A. Fedrizzi *et al.*, *Nature Phys.* **5**, 389 (2009) [2]

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S. Nauerth *et al.*, Nature Phot. 7, 382 (2013) [3]

From rooftops to satellites



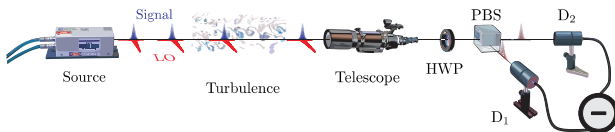
| [references](#)

J. Wang et al., Nature Phot. 7, 387 (2013) [4]

- ♣ Introduction
- ♠ Gaussian entanglement
- ♥ Higher-order moments
- ♦ Simulating atmospheric channels
- ♣ Summary and outlook

Basic example

Example: homodyne detection for atmosphere channels^{1,2}

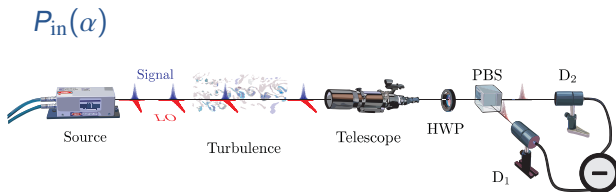


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- D. Elser et al., New J. Phys. **11**, 045014 (2009) [1]
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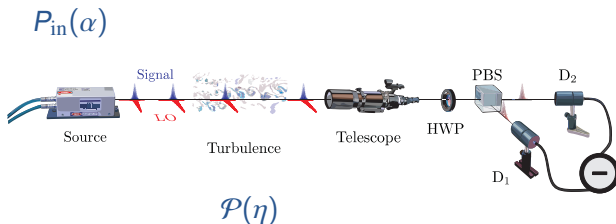


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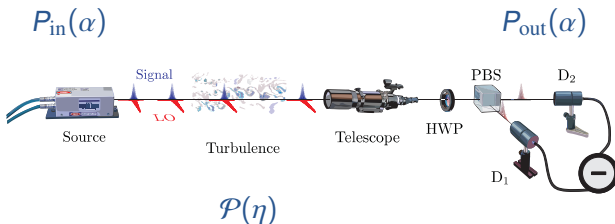


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Basic example

Example: homodyne detection for atmosphere channels^{1,2}



$$P_{out}(\alpha) = \int_0^1 d\eta \frac{1}{\eta} P_{in}\left(\frac{\alpha}{\sqrt{\eta}}\right) \mathcal{P}(\eta), \quad \eta = T^2 \in \mathbb{R}$$

| references

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Two main tasks

$$P_{\text{out}}(\alpha) = \int_0^1 d\eta \frac{1}{\eta} P_{\text{in}}\left(\frac{\alpha}{\sqrt{\eta}}\right) \mathcal{P}(\eta)$$



Martin Bohmann



Dr. Andrii Semenov



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$$P_{\text{out}}(\alpha) = \int_0^1 d\eta \frac{1}{\eta} P_{\text{in}}\left(\frac{\alpha}{\sqrt{\eta}}\right) \mathcal{P}(\eta)$$

Theory of the PDT



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Two main tasks

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Quantum effects in the atmosphere

Theory of the PDT



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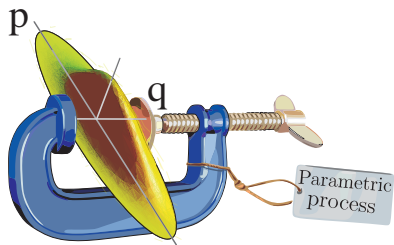


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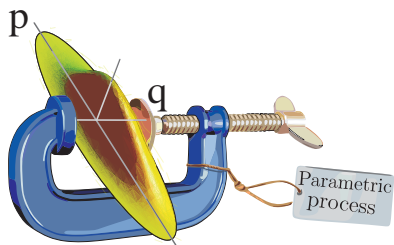
Quadrature squeezing



$$\langle \Delta \hat{q}^2 \rangle = \langle : \Delta \hat{q}^2 : \rangle + 1$$

$$\langle : \Delta \hat{q}^2 : \rangle < 1 \Rightarrow \text{squeezing}$$

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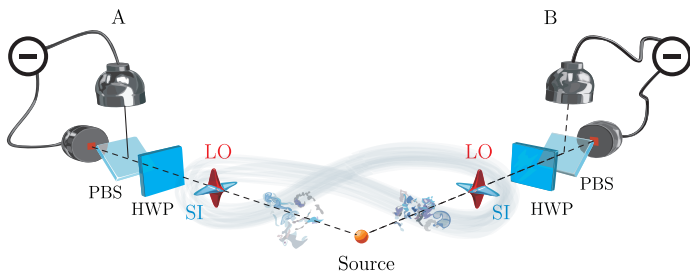
Input-output relation

$$\langle : \Delta \hat{q}^2 : \rangle_{\text{out}} = \underbrace{\langle T^2 \rangle}_{\text{standard attenuation}} \langle : \Delta \hat{q}^2 : \rangle_{\text{in}} + \underbrace{\langle \Delta T^2 \rangle \langle q \rangle_{\text{in}}^2}_{\text{turbulence}}$$

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Gaussian entanglement in the atmosphere

Basic setup:



Gaussian entanglement criterion

Simon criterion¹ which can be written as²

$$\mathcal{W} = \det V^{\text{PT}} < 0,$$

$$\text{with } V = \begin{pmatrix} \langle \Delta \hat{a}^\dagger \Delta \hat{a} \rangle & \langle \Delta \hat{a}^{\dagger 2} \rangle & \langle \Delta \hat{a}^\dagger \Delta \hat{b} \rangle & \langle \Delta \hat{a}^\dagger \Delta \hat{b}^\dagger \rangle \\ \langle \Delta \hat{a}^2 \rangle & \langle \Delta \hat{a} \Delta \hat{a}^\dagger \rangle & \langle \Delta \hat{a} \Delta \hat{b} \rangle & \langle \Delta \hat{a} \Delta \hat{b}^\dagger \rangle \\ \langle \Delta \hat{a} \Delta \hat{b}^\dagger \rangle & \langle \Delta \hat{a}^\dagger \Delta \hat{b}^\dagger \rangle & \langle \Delta \hat{b}^\dagger \Delta \hat{b} \rangle & \langle \Delta \hat{b}^{\dagger 2} \rangle \\ \langle \Delta \hat{a} \Delta \hat{b} \rangle & \langle \Delta \hat{a}^\dagger \Delta \hat{b} \rangle & \langle \Delta \hat{b}^2 \rangle & \langle \Delta \hat{b} \Delta \hat{b}^\dagger \rangle \end{pmatrix} = \begin{pmatrix} A & C^\dagger \\ C & B \end{pmatrix}$$

| references

- R. Simon, Phys. Rev. Lett. **84**, 2726 (2000) [1]
E. Shchukin and W. Vogel, Phys. Rev. Lett. **95**, 230502 (2005) [2]

Output condition¹

The fluctuating losses transform the covariance matrix V into

$$V \mapsto V_{\text{atm.}} = V_{\langle T_a^2 \rangle, \langle T_b^2 \rangle, \Gamma} + \begin{pmatrix} \vec{\mu}_a \vec{\mu}_a^\dagger & \Delta\Gamma \vec{\mu}_a \vec{\mu}_b^\dagger \\ \Delta\Gamma \vec{\mu}_b \vec{\mu}_a^\dagger & \vec{\mu}_b \vec{\mu}_b^\dagger \end{pmatrix}.$$

- displacement vectors $\vec{\mu}_a = \sqrt{\langle \Delta T_a^2 \rangle} (\langle \hat{a}^\dagger \rangle, \langle \hat{a} \rangle)^T$
- $V_{\text{atm.}}$ are reduced due to Γ : $C \mapsto \Gamma C$ and $\Gamma \in [0, 1]$
- Constant losses lead only to scaling $V_{\text{cl}} = T^2 V$

| **references**

M. B., A. A. Semenov, J. Sperling, and W. Vogel, Phys. Rev. A **94**, 010302(R) (2016). [1]

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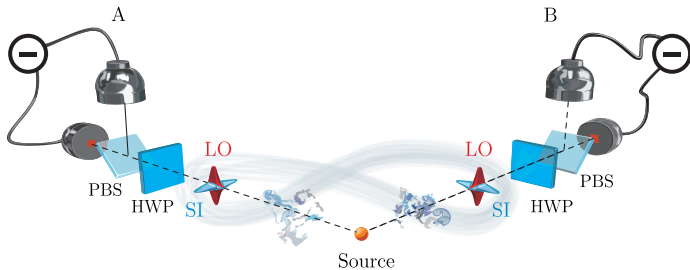
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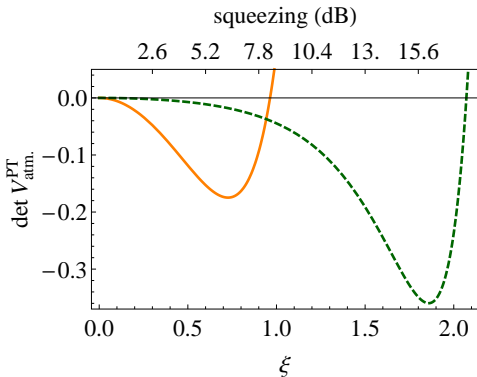
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M. B., A. A. Semenov, J. Sperling, and W. Vogel, Phys. Rev. A **94**, 010302(R) (2016). [1]



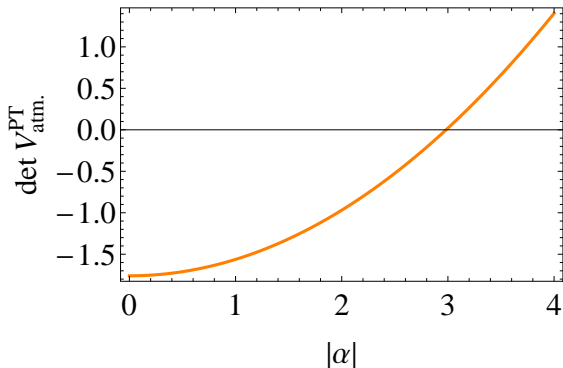
Now we consider uncorrelated loss ($T_a \neq T_b$)

Dependence on squeezing strength

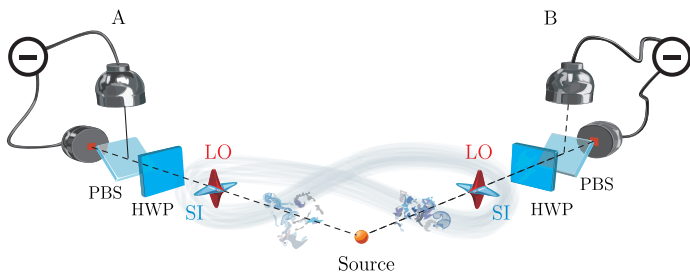


- Two-mode squeezed vacuum state in 144 km and 1.6 km free-space channel
- Strong squeezing leads to disentanglement

Dependence on coherent displacement



- Displaced two-mode squeezed vacuum state
- Displacement can lead to disentanglement

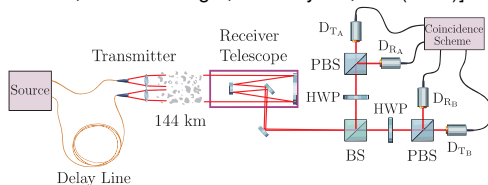


Now we consider correlated loss ($T_a = T_b$)

Correlated channels: $T_a = T_b$

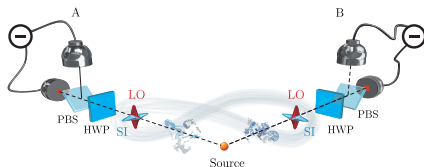
Copropagation

Discrete variables: A. Fedrizzi, R. Ursin, T. Herbst, M. Nespoli, R. Prevedel, T. Scheidl, F. Tiefenbacher, T. Jennewein, and A. Zeilinger, *Nat. Phys.* **5**, 389 (2009)]



Correlated channels: $T_a = T_b$

Adaptive correlations



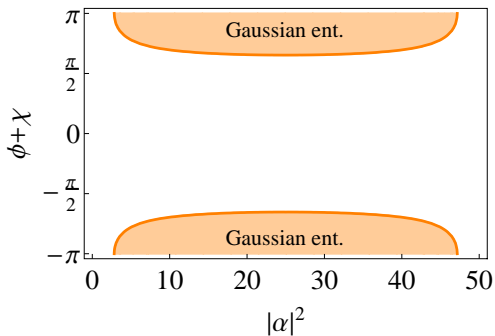
1. Monitor both transmission coefficients T_a and T_b
2. Share this information via classical communication
3. Attenuate one channel to the level of the other $\Rightarrow T_a = T_b$
4. Result: corresponds to deterministic loss

Most Gaussian entanglement is preserved²

| references

- M. B., A. A. Semenov, J. Sperling, and W. Vogel, Phys. Rev. A **94**, 010302(R) (2016). [1]
F. A. S. Barbosa *et al.*, Phys. Rev. A **84**, 052330 (2011). [2]

Dependence on coherent displacement



- Displaced two-mode squeezed vacuum state:
 $\langle \hat{a} \rangle = |\alpha|e^{i\phi}$ and $\langle \hat{b} \rangle = |\beta|e^{i\chi}$
- Plot: $|\alpha|^2 + |\beta|^2 = 50$, 144 km atmospheric channel
- Gaussian entanglement depends on displacement direction

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Moment criteria

Method: matrix of moments $M_{(\vec{p},\vec{q}),(\vec{r},\vec{s})} = \langle [\hat{a}^{\dagger\vec{p}} \hat{a}^{\vec{q}}]^{\dagger} [\hat{a}^{\dagger\vec{r}} \hat{a}^{\vec{s}}] \rangle$ ^{1,2}

$$N_{(\vec{p},\vec{q}),(\vec{r},\vec{s})} < 0 \xleftarrow[\text{ordering}]{\text{normal}} M_{(\vec{p},\vec{q}),(\vec{r},\vec{s})} \xrightarrow[\text{transposition}]{\text{partial}} M_{(\vec{p},\vec{q}),(\vec{r},\vec{s})}^{\text{PT}} < 0$$

nonclassicality

PT entanglement

| references

- E. Shchukin, Th. Richter, and W. Vogel, Phys. Rev. A **71**, 011802(R) (2005) [1]
- E. Shchukin and W. Vogel, Phys. Rev. Lett. **95**, 230502 (2005) [2]
- M. Bohmann, J. Sperling, A. A. Semenov, and W. Vogel, Phys. Rev. A **95**, 012324 (2017) [2]

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Derivation of the input-output relation³

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Amplitude squeezing

- Criterion for amplitude squeezing:

$$d = \begin{vmatrix} 1 & \langle \hat{a}^\dagger \rangle & \langle \hat{a} \rangle \\ \langle \hat{a} \rangle & \langle \hat{a}^\dagger \hat{a} \rangle & \langle \hat{a}^2 \rangle \\ \langle \hat{a}^\dagger \rangle & \langle \hat{a}^{\dagger 2} \rangle & \langle \hat{a}^\dagger \hat{a} \rangle \end{vmatrix} = \underbrace{\begin{vmatrix} \langle \Delta \hat{a}^\dagger \Delta \hat{a} \rangle & \langle (\Delta \hat{a})^2 \rangle \\ \langle (\Delta \hat{a}^\dagger)^2 \rangle & \langle \Delta \hat{a}^\dagger \Delta \hat{a} \rangle \end{vmatrix}}_A < 0$$

- Output condition:

$$d^{\text{out}} = \langle T^2 \rangle^2 \left[d + \Gamma (\langle \hat{a}^\dagger \rangle, \langle \hat{a} \rangle) A \begin{pmatrix} \langle \hat{a} \rangle \\ \langle \hat{a}^\dagger \rangle \end{pmatrix} \right] < 0$$

$$\text{with } \Gamma = 1 - \frac{\langle T \rangle^2}{\langle T^2 \rangle}$$

⇒ strong dependency on the coherent displacement

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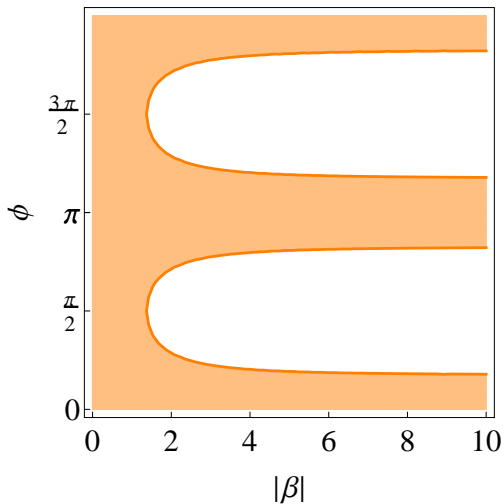
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Amplitude squeezing



Higher-order entanglement

- Example of a higher-order criterion:

$$S = \left| \begin{array}{cc} 1 & \langle \hat{a} \hat{b}^\dagger \rangle \\ \langle \hat{a}^\dagger \hat{b} \rangle & \langle \hat{a}^\dagger \hat{a} \hat{b}^\dagger \hat{b} \rangle \end{array} \right|_{\text{ent.}} < 0$$

- Output condition:

$$S^{\text{out}} = \langle T_a^2 T_b^2 \rangle [S + \Gamma_{ab} |\langle \hat{a} \hat{b}^\dagger \rangle|^2]$$

with $\Gamma_{ab} = \frac{\langle T_a^2 T_b^2 \rangle - \langle T_a T_b \rangle^2}{\langle T_a^2 T_b^2 \rangle}$ and $\Gamma_{ab} \in [0, 1]$

- Is non-Gaussian entanglement more or less robust?

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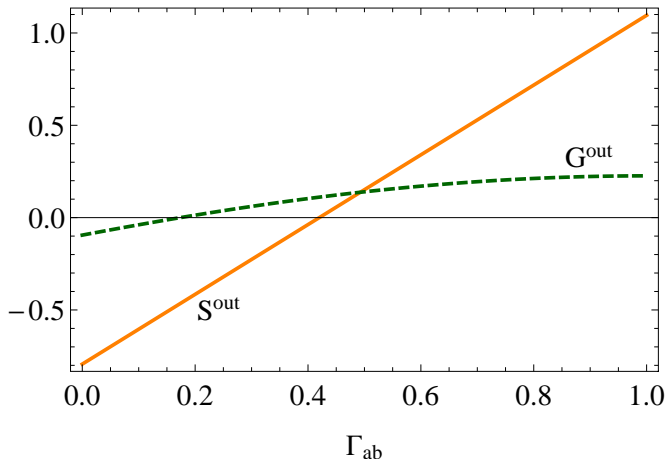
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Higher-order entanglement

State: $|\Psi^-\rangle = \mathcal{N}(\alpha)(|\alpha\rangle \otimes |\alpha\rangle - |-\alpha\rangle \otimes |-\alpha\rangle)$



Outline

- ♣ Introduction
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- ♥ Higher-order moments
- ♦ **Simulating atmospheric channels**
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Idea¹

- **Aim:** testing atmospheric channels in the lab
- Measure with different constant attentions η_k
- Allows to approximate the atmospheric transmission distribution

$$\tilde{\mathcal{P}}(Y = \eta_k) = \mathcal{P}(\eta_k) \left[\sum_{j=0}^n \mathcal{P}(\eta_j) \right]^{-1}.$$

- Measurement results

$$x^{\text{atm.}} = \sum_{j=0}^n \tilde{\mathcal{P}}(\eta_j) x_j.$$

| **references**

M. Bohmann, R. Kruse, J. Sperling, C. Silberhorn, and W. Vogel, Phys. Rev. A **95**, 063801 (2017). [1]

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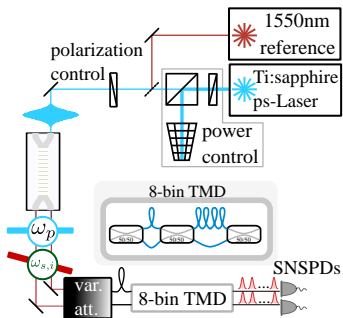
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M. Bohmann, R. Kruse, J. Sperling, C. Silberhorn, and W. Vogel, Phys. Rev. A **95**, 063801 (2017). [1]

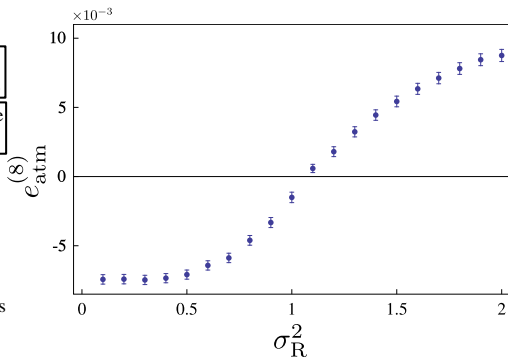
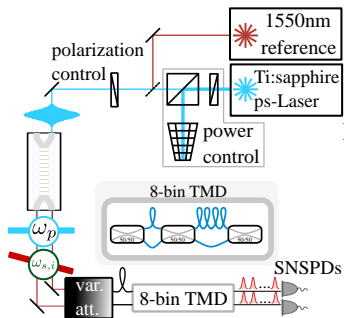
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M. Bohmann, R. Kruse, J. Sperling, C. Silberhorn, and W. Vogel, Phys. Rev. A **95**, 063801 (2017). [1]

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- ♣ Introduction
- ♠ Gaussian entanglement
- ♥ Higher-order moments
- ♦ Simulating atmospheric channels
- ♣ **Summary and outlook**

Conclusion

- Full treatment of Gaussian entanglement in the atmosphere
- Dependence on squeezing and coherent displacements
- Preserving Gaussian entanglement with adoptive channel correlation
- Generalization to nonclassicality and PT entanglement testing
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Outlook

How do our results relate to

- Quantum key distribution¹
- Entanglement and squeezing distillation²
- Quantum teleportation³
- Sensing the atmosphere

| references

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Thank you for your attention!